Computer Graphics III Spherical integrals, Light & Radiometry – Exercises

Jaroslav Křivánek, MFF UK

Jaroslav.Krivanek@mff.cuni.cz

Surface area of a (subset of a) sphere

- Calculate the surface area of a unit sphere.
- Calculate the surface area of a spherical cap delimited by the angle θ_0 measured from the north pole.
- Calculate the surface area of a spherical wedge with angle ϕ_0 .

Solid angle

- What is the solid angle under which we observe an (infinite) plane from a point outside of the plane?
- Calculate the solid angle under which we observe a sphere with radius *R*, the center of which is at the distance *D* from the observer.

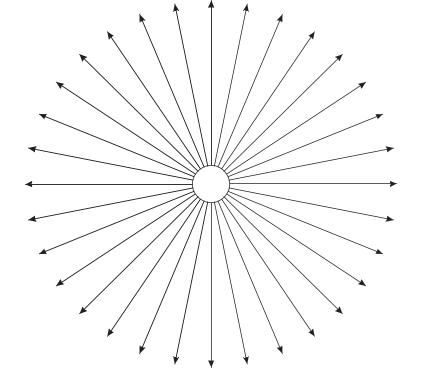
Isotropic point light

• **Q:** What is the emitted power (flux) of an isotropic point light source with intensity that is a constant *I* in all directions?

Isotropic point light

• **A:** Total flux:

$$\Phi = \int_{\Omega} I(\omega) \, d\omega = \begin{vmatrix} substitute : \\ d\omega = \sin\theta \, d\theta \, d\varphi \end{vmatrix}$$
$$= I \int_{\varphi=0}^{2\pi} \int_{\theta=0}^{\pi} \sin\theta \, d\theta \, d\varphi$$
$$= I 2\pi [-\cos\theta]_{0}^{\pi}$$
$$= 4\pi I$$

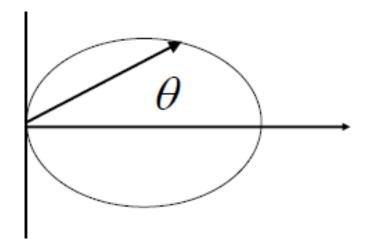


$$I = \frac{\Phi}{4\pi}$$

Cosine spot light

 What is the power (flux) of a point source with radiant intensity given by:

 $I(\omega) = I_0 \max\{0, \omega \cdot \vec{d}\}^s$



Spotlight with linear angular fall-off

• What is the power (flux) of a point light source with radiant intensity given by:

$$I(\theta, \phi) = \begin{cases} I_0 & \theta \le \alpha \\ I_0 \frac{\beta - \theta}{\beta - \alpha} & \alpha < \theta < \beta \\ 0 & \theta \ge \beta \end{cases}$$

Constant part

$$\Phi_1 = \int_0^{2\pi} \int_0^{\alpha} I_0 \sin \theta d\theta d\phi = I_0 2\pi (1 - \cos \alpha).$$

Linear part

$$\Phi_2 = \int_0^{2\pi} \int_\alpha^\beta I_0 \frac{\beta - \theta}{\beta - \alpha} \sin \theta d\theta d\phi = I_0 \frac{2\pi}{\beta - \alpha} \int_\alpha^\beta (\beta - \theta) \sin \theta d\theta \tag{1}$$

The last integral is the sum of the following two integrals:

$$\int_{\alpha}^{\beta} \beta \sin \theta d\theta = \beta \cos \alpha - \beta \cos \beta \tag{2}$$

$$-\int_{\alpha}^{\beta} \theta \sin \theta d\theta = \left| \sin \theta - \theta \cos \theta \right|_{\beta}^{\alpha} = \sin \alpha - \alpha \cos \alpha - \sin \beta + \beta \cos \beta$$
(3)

Plugging (2) and (3) into (1) and rearranging, we get

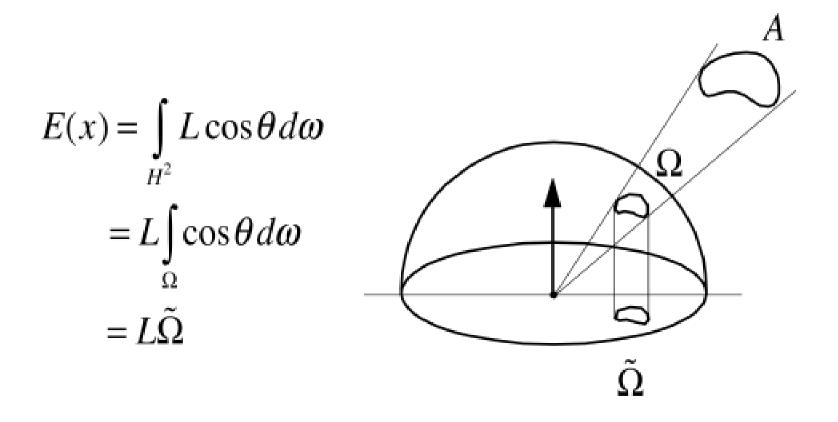
$$\Phi_2 = I_0 \frac{2\pi}{\beta - \alpha} \left[(\beta - \alpha) \cos \alpha + \sin \alpha - \sin \beta \right] = I_0 2\pi \left[\cos \alpha - \frac{\sin \beta - \sin \alpha}{\beta - \alpha} \right]. \tag{4}$$

Total flux

$$\Phi = \Phi_1 + \Phi_2 = I_0 2\pi \left[1 - \frac{\sin\beta - \sin\alpha}{\beta - \alpha} \right]$$
CG III (NPGR010) - J. Křivánek 2016
(5)

Irradiance due to a Lambertian light source

What is the irradiance *E*(**x**) at point **x** due to a uniform Lambertian area source observed from point **x** under the solid angle Ω?



Pat Hanrahan, 2009

How dark are outdoor shadows?

- luminance arriving on a surface from a full (overhead) sun is 300,000 × luminance arriving from the blue sky, but the sun occupies only a small fraction of the sky
- illuminance on a sunny day = 80% from the sun + 20% from blue sky, so shadows are 1/5 as bright as lit areas (2.3 f/stops)

mean = 7 m

mean = 27

Based in these hints, calculate the solid angle under which we observe the Sun. (We assume the Sun is at the zenith.)

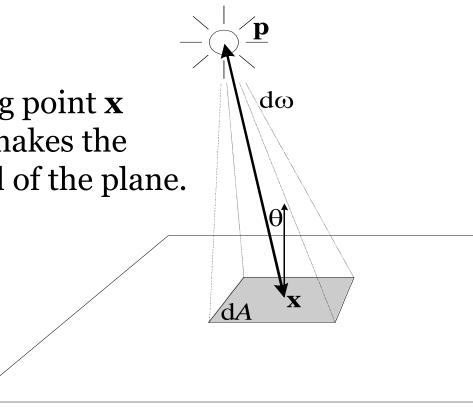
(Marc Levoy)

19



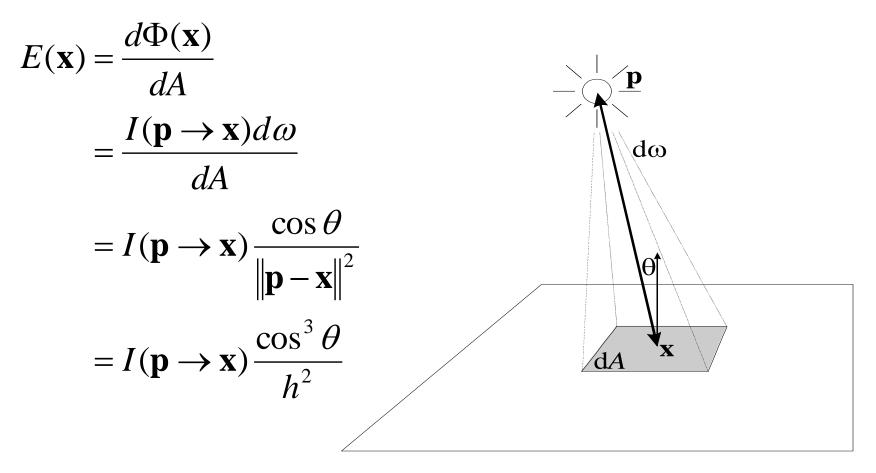
Irradiance due to a point source

- What is the irradiance at point **x** on a plane due to a point source with intensity *I*(ω) placed at the height *h* above the plane.
- The segment connecting point x to the light position p makes the angle θ with the normal of the plane.



Irradiance due to a point source

Irradiance of a point on a plane lit by a point source:



CG III (NPGR010) - J. Křivánek 2016

Area light sources

- Emission of an area light source is fully described by the emitted radiance *L_e*(**x**,ω) for all positions on the source **x** and all directions ω.
- The total emitted power (flux) is given by an integral of L_e(**x**,ω) over the surface of the light source and all directions.

$$\Phi = \int_{A} \int_{H(\mathbf{x})} L_e(\mathbf{x},\omega) \cos\theta \,\mathrm{d}\omega \,\mathrm{d}A$$

Diffuse (Lambertian) light source

What is the relationship between the emitted radiant exitance (radiosity) B_e(x) and emitted radiance L_e(x, ω) for a Lambertian area light source?

Lambertian source =

emitted radiance does not depend on the direction $\boldsymbol{\omega}$

 $L_e(\mathbf{x}, \omega) = L_e(\mathbf{x}).$

Diffuse (Lambertian) light source

- $L_e(\mathbf{x}, \omega)$ is constant in ω
- Radiosity: $B_e(\mathbf{x}) = \pi L_e(\mathbf{x})$

$$B_e(\mathbf{x}) = \int_{H(\mathbf{x})} L_e(\mathbf{x}, \omega) \cos \theta \, \mathrm{d}\omega$$
$$= L_e(\mathbf{x}) \int_{H(\mathbf{x})} \cos \theta \, \mathrm{d}\omega$$
$$= \pi L_e(\mathbf{x})$$

Uniform Lambertian light source

- What is the total emitted power (flux) Φ of a **uniform** Lambertian area light source which emits radiance L_e
 - Uniform source radiance does not depend on the position, $L_e(\mathbf{x}, \omega) = L_e = \text{const.}$

Uniform Lambertian light source

• $L_e(\mathbf{x}, \boldsymbol{\omega})$ is constant in \mathbf{x} and $\boldsymbol{\omega}$

$$\Phi_e = \mathbf{A} \mathbf{B}_e = \pi \mathbf{A} L_e$$